# **Convective flow velocity measurements** in a rapidly rotating sphere

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### Oersted 2000



Dipolar magnetic field

### Earth's magnetic field at the Core-Mantle Boundary





### Oersted 2000, at the CMB

#### Velocity Field at the CMB

### The Earth's Core





Structure and composition of the Earth's Core

Flow at the onset of convection in a rapidly rotating sphere

# Our experimental approach of the dynamical flow of the Earth's Core







Vortex of Water, Gal- Thermal convection in<br/>lium, SodiumGal- Thermal convection in<br/>GalliumThermal convection in<br/>Gallium

# **MEASUREMENTS OF** V(t)





Doppler apparatus

## Velocity measurements in vortices of water, gallium, sodium



# Thermal convection in a rapidly rotating sphere: flow visualisation



 $\begin{array}{c} \textbf{Coriolis Force} \Longrightarrow \\ - \textbf{Geostrophic flow} \end{array}$ 



### Velocity measurements in a rapidly rotating sphere of water: radial velocities

### Velocity measurements in a rapidly rotating sphere of water : different dynamical regimes at fixed rotation rate



Evolution of time-depth radial patterns with Ra at  $E = 9.7 \ 10^{-6}$ 

# Velocity measurements in a rapidly rotating sphere of water : different regimes at fixed $Ra/Ra_c$







## Comparison of profiles obtained in Water and liquid Gallium.



## Non-dimensional numbers for the convection flow.

Number	Name	Meaning	Water	Gallium	Earth's core
$Ra = \frac{\alpha \Delta T g_D D^3}{\kappa \nu}$	Rayleigh number	buoyancy viscosity	$10^7 - 10^9$	$10^7 - 10^8$	?
$E = \frac{\nu}{\Omega D^2}$	Ekman number	$\frac{VISCOSILY}{Coriolis}$	$10^{-5} - 10^{-6}$	$10^{-6} - 10^{-7}$	$10^{-15} - 10^{-13}$
$P = \frac{\nu}{\kappa}$	Prandtl number	$\frac{\text{viscous diffusivity}}{\text{thermal diffusivity}}$	7	0.025	0.1 - 10



Velocity fields are accessible in MHD experiments using the Doppler ultrasonic technique.

We have established power scaling laws for the explored parameter regime (Pr, E, Ra).

If the results are extrapolated to the core, they give structures which are too thin. Influence of the magnetic field...? FUTURE:

Magneto-convection in a rapidly rotating sphere of gallium.

Moderate  $R_m$  experiments in Grenoble with liquid sodium under rotation and strong imposed magnetic field in a spherical geometry.





Let  $u, \delta, T$  be scales for equatorial velocity, cell size and temperature perturbation. From convective heat flow over a sphere we get

$$T = 4\pi \frac{Nu}{uP^2},$$

Inertial balance: we equate Vortex stretching, Inertia, and buoyancy.

$$u = \left[ (4\pi Nu)^{2/5} \left(\frac{2}{L} \frac{dL}{dr}\right)^{-1/5} \right] \left(\frac{Ra}{P^2}\right)^{2/5} E^{1/5}$$
$$\delta = \left[ (4\pi Nu)^{1/5} \left(\frac{2}{L} \frac{dL}{dr}\right)^{-3/5} \right] \left(\frac{Ra}{P^2}\right)^{1/5} E^{3/5}$$
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Vorticity  $\boldsymbol{\omega}$  reduces to its z component  $\boldsymbol{\omega}$ , and the resulting equation writes:

$$\frac{\partial\omega}{\partial t} + \mathbf{u} \cdot \nabla\omega - \beta \mathbf{u} \cdot \mathbf{e}_r = \nabla^2 \omega + \frac{Ra}{r} \frac{\partial T}{\partial \theta},\tag{1}$$

with

$$\beta = \frac{2}{L} \frac{\mathrm{d}L}{\mathrm{d}r} E^{-1},$$

and

$$L = \sqrt{1 - \left(\frac{r}{r_e}\right)^2}.$$

L is half the height of a vertical fluid column between the upper and lower boundaries.



Time averaged zonal flow speed measured along radius

**Radius-averaged zonal speed against** Ra

 $E = 9.8 \ 10^{-6}$